

New Challenges in the Kinetic Theory of Complex Fluids



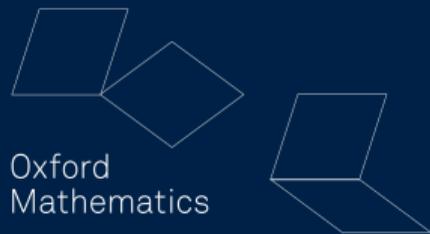
Mathematical
Institute

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A. Medaglia*, G. Russo†

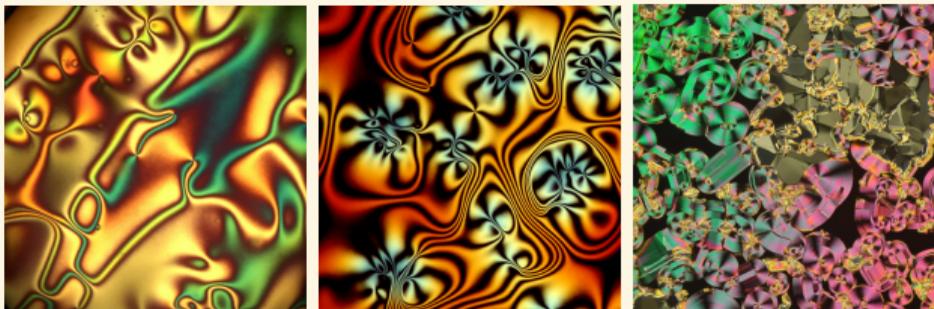
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- ▶ **Liquid crystals**, here depicted in nematic and smectic phases.
- ▶ **Ferrofluids**, i.e. a colloidal suspension made of nanoscale ferromagnetic or ferrimagnetic particles.
- ▶ **Gas saturated magma melts** and other fluids with non-diffusive bubbles.

ORDER PARAMETER MANIFOLD

1

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Order Parameter Manifold

We say that the tuple $(\mathcal{M}, \mathcal{A})$ is an order parameter manifold if \mathcal{M} is a smooth manifold with a fixed parametrization, and \mathcal{A} is a Lie group action of $SO(d)$ on \mathcal{M} , i.e. the map \mathcal{A} is smooth enough to be differentiable.



Continua with microstructure,
(G. Capriz),
Differential geometry and continuum mechanics, (G. Capriz,
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Furthermore, we say that a field $\nu : \mathbb{E}^d \rightarrow \mathcal{M}$ is an order parameter field if $\forall \underline{c} \in \mathbb{R}^d$ and $\forall \underline{Q} \in SO(d)$ we have

$$\nu(\underline{Q}\mathbf{x} + \underline{c}) = \mathcal{A}(\underline{Q}, \nu(\mathbf{x})), \quad \forall \mathbf{x} \in \mathbb{E}^d.$$



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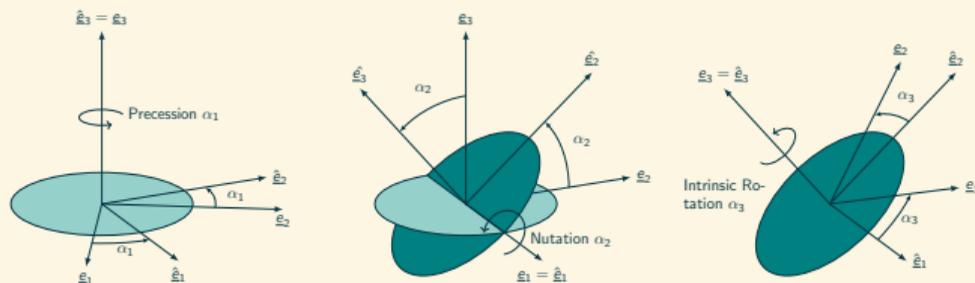
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- ▶ We need to understand what manifold \mathcal{M} captures the nature of the order parameters.
- ▶ We need to understand the action of rotations on the manifold \mathcal{M} .

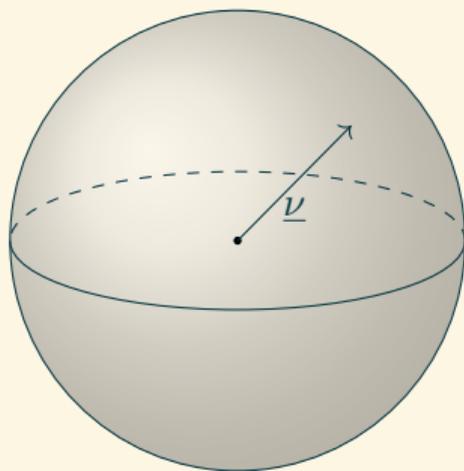
AN EXAMPLE: NEMATIC LIQUID CRYSTALS



Variational Theories for Liquid Crystals, (E. Virga),
The Physics of Liquid Crystals,
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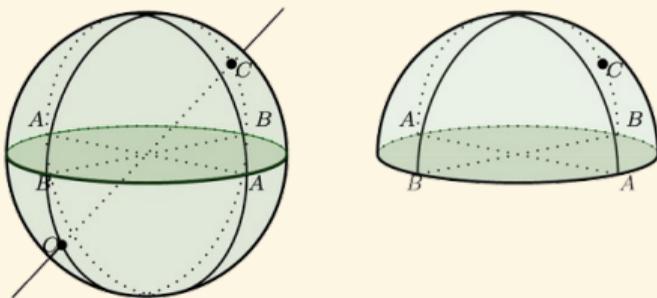
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- ▶ We can represent the state of a calamitic molecule using the set of Euler angles θ, ϕ, ψ .
- ▶ We can also represent the state of a calamitic molecule using a director field $\underline{v} \in \mathbb{S}^2$.
- ▶ For head-tail symmetric calamitic molecules, we can use \mathbb{RP}^2 .

EMBEDDING RESULTS

Embedding theorems

- ▶ Any compact orientable 2-manifold can be embedded in \mathbb{R}^3 .



Curves and Surfaces, (M. Abate,
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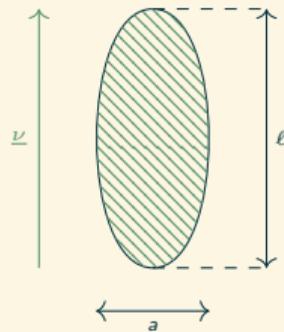
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THE MICROSCOPIC WORLD

2

LAGRANGIAN MECHANICS OF THE CONSTITUENTS

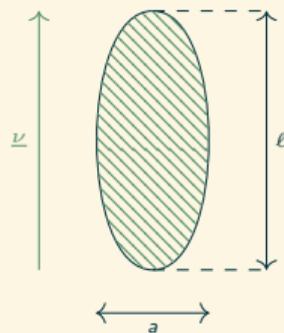
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$$\mathcal{L}_i := \frac{1}{2} m_1 (\dot{\underline{x}}_i \cdot \dot{\underline{x}}_i) + \frac{1}{2} \dot{\underline{\nu}}_i \cdot \underline{\underline{\Omega}}_i(\nu_i) \dot{\underline{\nu}}_i.$$



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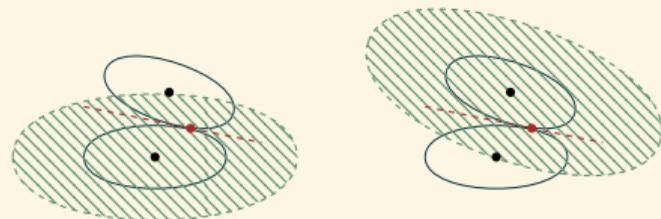
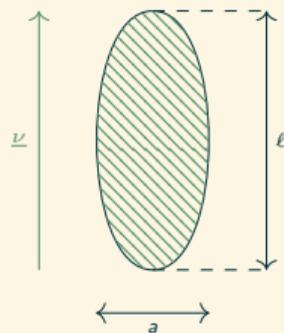
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We assume the interaction between the constituents is given by a potential $\mathcal{W}(\mathbf{x}_i - \mathbf{x}_j, \nu_i, \nu_j)$, i.e.

$$\mathcal{L}_{i,j} = \mathcal{L}_i(\mathbf{x}_i, \underline{\Xi}_i) + \mathcal{L}_j(\mathbf{x}_j, \underline{\Xi}_j) + \mathcal{W}(\mathbf{x}_i - \mathbf{x}_j, \nu_i, \nu_j),$$

where $\underline{\Xi}_i := (\underline{v}_i, \nu_i, \underline{\dot{\nu}}_i)$.



NOETHER'S THEOREM: SYMMETRIES AND CONSERVATION LAWS

Noether's theorem

If a Lagrangian \mathcal{L} is invariant under a group action with infinitesimal generators G then

$$\frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{q}_{1,2}} \cdot G \right) = 0, \quad q_{1,2} = (\mathbf{x}_1, \mathbf{x}_2, \nu_1, \nu_2).$$

In other words for any physical symmetry of the system, there is a conserved quantity.



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- ▶ The Lagrangian \mathcal{L} is invariant under translations, i.e. the linear momentum is conserved.
- ▶ The Lagrangian \mathcal{L} is independent of time and the kinetic energy is a homogeneous quadratic form of the conjugate moments, i.e. the energy is conserved.

NOETHER'S THEOREM: SYMMETRIES AND CONSERVATION LAWS

Infinitesimal Generator of \mathcal{A}

For fixed $\nu \in \mathcal{M}$, the orbit map

$$\mathcal{A}_\nu : SO(3) \rightarrow SO(3)\nu, \quad \underline{Q} \mapsto \mathcal{A}(\underline{Q}, \nu),$$

is differentiable at the identity.

We will denote by $A_\nu : SO(3) \rightarrow T_\nu\mathcal{M}$ the differential of \mathcal{A}_ν at the identity.

Composing the canonical isomorphism $\mathbb{R}^3 \rightarrow SO(3)$ with the differential of the orbit map we obtain a map $A_\nu : \mathbb{R}^3 \rightarrow T_\nu\mathcal{M}$.



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Assuming that the Lagrangian \mathcal{L} is frame-indifferent, i.e. invariant under the action of $\text{SO}(3)$, we have:

$$G = (\underline{r} \times \mathbf{x}, \underline{r} \times \underline{x}, A_\nu \mathbf{r}, A_\nu \underline{r}),$$

where \underline{r} is the rotation axis. Thus, the angular momentum is conserved.

AN EXAMPLE: ANGULAR MOMENTUM NEMATIC LIQUID CRYSTALS

For segment like molecules the classical we have $\underline{\underline{\Omega}}(\underline{\nu}) = I$, where I is the Identity. Thus, Noether's theorem implies the conservation of the following quantity:

$$m_1 \mathbf{x}_1 \times \underline{p}_1 + \nu \times \underline{\dot{\nu}}_1 + m_2 \mathbf{x}_1 \times \underline{p}_2 + \nu \times \underline{\dot{\nu}}_2.$$

Let $\underline{\omega}$ be the angular velocity of the segment, using the triple cross product together with the well-known property of segment like rigid bodies that $\underline{\dot{\nu}}_i = \underline{\omega} \times \underline{\nu}_i$ we can rewrite one term of the previous expression as

$$\underline{\nu}_i \times \underline{\dot{\nu}}_i = \underline{\nu}_i \times \underline{\omega}_i \times \underline{\nu}_i = (\underline{\nu}_i \cdot \underline{\nu}_i) \underline{\omega} - (\underline{\nu}_i \cdot \underline{\omega}_i) \underline{\nu}_i = \underline{\omega}_i - (\underline{\nu}_i \cdot \underline{\omega}_i) \underline{\nu}_i = \mathbb{I}_i \underline{\omega},$$

where used the fact that the inertia tensor of a segment is $\mathbb{I}_i := I - \underline{\nu}_i \otimes \underline{\nu}_i$. Therefore, we retrieved the classical definition of angular momentum, i.e.

$$\mathbf{x}_1 \times \underline{p}_1 + \mathbb{I}_1 \underline{\omega}_1 + \mathbf{x}_2 \times \underline{p}_2 + \mathbb{I}_2 \underline{\omega}_2,$$

BBGKY HIERARCHY

3

HAMILTONIAN MECHANICS OF THE CONSTITUENTS

We introduce the Hamiltonian formalism associated to the Lagrangian \mathcal{L} introduced in the previous section. As usual, we introduce the conjugate momenta to the generalised coordinates, i.e.

$$\underline{p}_i := \frac{\partial \mathcal{L}}{\partial \dot{\mathbf{x}}_i} = m \dot{\mathbf{x}}_i, \quad \underline{s}_i := \frac{\partial \mathcal{L}}{\partial \dot{\nu}_i} = \underline{\underline{\Omega}}(\nu) \dot{\nu}_i.$$

We then introduce the Hamiltonian \mathcal{H} of the full system of N constituents, only interacting in pairs, as

$$\mathcal{H} := \sum_{i=1}^N \frac{1}{2m} \underline{p}_i \cdot \underline{p}_i + \frac{1}{2} \underline{s}_i \cdot \underline{\underline{\Omega}}(\nu)^{-1} \underline{s}_i + \sum_{1 \leq i < j \leq N} \mathcal{W}(\mathbf{x}_i - \mathbf{x}_j, \nu_i, \nu_j).$$

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The Legendre transform of the Lagrangian \mathcal{L} is always well-defined, assuming $\underline{\underline{\Omega}}(\nu)$ is symmetric and positive definite for all $\nu \in \mathcal{M}$.

BOGOLIUBOV–BORN–GREEN–KIRKWOOD–YVON HIERARCHY



An Introduction to the Theory of the Boltzmann Equation, (S. Harris),
Statistical Physics of Particles, (M. Kardar),
Statistical Mechanics, 2nd Edition (K. Huang).

Let f_s denote the normalised s -particle distribution function. We obtain the following expression for the BBGKY hierarchy,

$$\begin{aligned} \frac{\partial f_s}{\partial t} + \{\pi_s, \mathcal{H}_s\} &= \int \sum_{i=1}^s \frac{\partial f_{s+1}}{\partial \underline{p}_i} \cdot \frac{\partial \mathcal{W}(\mathbf{x}_i - \mathbf{x}_{s+1}, \nu_i, \nu_{s+1})}{\partial \mathbf{x}_i} d\Gamma_{s+1} \\ &+ \int \sum_{i=1}^s \frac{\partial f_{s+1}}{\partial \underline{c}_i} \cdot \frac{\partial \mathcal{W}(\mathbf{x}_i - \mathbf{x}_{s+1}, \nu_i, \nu_{s+1})}{\partial \nu_i} d\Gamma_{s+1}, \end{aligned}$$

where $\mathcal{H}_s = \left(\sum_{i=1}^s \frac{|p_i|^2}{2m} + \frac{1}{2} \underline{c}_i \cdot \underline{\underline{\Omega}}(\nu)^{-1} \underline{c}_i \right) + \sum_{1 \leq i < j \leq s} \mathcal{W}(|\mathbf{x}_i - \mathbf{x}_j|, \nu_i, \nu_j)$.

BOGOLIUBOV–BORN–GREEN–KIRKWOOD–YVON HIERARCHY

The first two terms of the BBGKY hierarchy, under the assumption that there are no three-body interactions, amount to

$$\begin{aligned}
 & \frac{\partial f_1}{\partial t} + \frac{p_1}{m} \cdot \frac{\partial f_1}{\partial \mathbf{x}_1} + \underline{\underline{\Omega}}(\nu_1)^{-1} \underline{\underline{\zeta}}_1 \frac{\partial f_1}{\partial \nu_1} = \\
 & + \int \frac{\partial \mathcal{W}(\mathbf{x}_1 - \mathbf{x}_2, \nu_1, \nu_2)}{\partial \mathbf{x}_1} \left(\frac{\partial f_2}{\partial \underline{p}_1} - \frac{\partial f_2}{\partial \underline{p}_2} \right) \\
 & + \int \frac{\partial \mathcal{W}(\mathbf{x}_1 - \mathbf{x}_2, \nu_1, \nu_2)}{\partial \nu_1} \left(\frac{\partial f_2}{\partial \underline{\zeta}_1} - \frac{\partial f_2}{\partial \underline{\zeta}_2} \right)
 \end{aligned}$$

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 & + \frac{p_2}{m} \cdot \frac{\partial f_2}{\partial \mathbf{x}_2} + \underline{\underline{\Omega}}(\nu_2)^{-1} \underline{\underline{\zeta}}_2 \cdot \frac{\partial f_2}{\partial \nu_2} \\
 & - \frac{\partial \mathcal{W}(\mathbf{x}_1 - \mathbf{x}_2, \nu_1, \nu_2)}{\partial \mathbf{x}_1} \left(\frac{\partial f_2}{\partial \underline{p}_1} - \frac{\partial f_2}{\partial \underline{p}_2} \right) \\
 & - \frac{\partial \mathcal{W}(\mathbf{x}_1 - \mathbf{x}_2, \nu_1, \nu_2)}{\partial \nu_1} \frac{\partial f_2}{\partial \underline{\zeta}_1} = 0
 \end{aligned}$$

BOGOLIUBOV–BORN–GREEN–KIRKWOOD–YVON HIERARCHY

To highlight the same timescale separation in the second term of the hierarchy we introduce fast and slow varying coordinates, i.e.

$$\mathbf{x} = \mathbf{x}_2 - \mathbf{x}_1, \quad \mathbf{X} = \frac{1}{2}(\mathbf{x}_2 + \mathbf{x}_1).$$

We then boxed the terms that are quickly varying in the second equation of the BBGKY hierarchy, i.e.

$$\frac{\partial f_2}{\partial t} + \frac{1}{2} \frac{\underline{p}_2 + \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \underline{X}} + \underline{\underline{\Omega}}(\nu_1)^{-1} \underline{\underline{\zeta}}_1 \cdot \frac{\partial f_2}{\partial \nu_1} + \underline{\underline{\Omega}}(\nu_2)^{-1} \underline{\underline{\zeta}}_2 \cdot \frac{\partial f_2}{\partial \nu_2} + \frac{\underline{p}_2 - \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \mathbf{x}}$$

$$- \frac{\partial \mathcal{W}(\mathbf{x}_1 - \mathbf{x}_2, \nu_1, \nu_2)}{\partial \mathbf{x}_1} \cdot \left(\frac{\partial f_2}{\partial \underline{p}_1} - \frac{\partial f_2}{\partial \underline{p}_2} \right) - \frac{\partial \mathcal{W}(|\mathbf{x}_1 - \mathbf{x}_2|, \nu_1, \nu_2)}{\partial \nu_1} \cdot \frac{\partial f_2}{\partial \underline{\zeta}_1} = 0$$

EMBEDDED BOGOLIUBOV–BORN–GREEN–KIRKWOOD–YVON HIERARCHY

Using the embedding results previously discussed, we can use the fast and slow varying coordinates also for the order parameters, i.e.

$$\underline{n} = \underline{\nu}_2 - \underline{\nu}_1, \quad \underline{N} = \frac{1}{2} (\underline{\nu}_2 + \underline{\nu}_1).$$

We then introduce $\underline{A} = \frac{1}{2} (\underline{\Omega}_2(\underline{\nu}_2)^{-1} \underline{\zeta}_1 + \underline{\Omega}_1(\underline{\nu}_1)^{-1} \underline{\zeta}_2)$, $\underline{B} = (\underline{\Omega}_2(\underline{\nu}_2)^{-1} \underline{\zeta}_2 - \underline{\Omega}_1(\underline{\nu}_1)^{-1} \underline{\zeta}_1)$, i.e.

$$\begin{aligned} & \frac{\partial f_2}{\partial t} + \frac{1}{2} \frac{\underline{p}_2 + \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \underline{X}} + \underline{A} \cdot \frac{\partial f_2}{\partial \underline{N}} + \boxed{\underline{B} \cdot \frac{\partial f_2}{\partial \underline{n}}} + \boxed{\frac{\underline{p}_2 - \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \underline{x}}} \\ & - \boxed{\frac{\partial \mathcal{W}(\underline{x}_1 - \underline{x}_2, \nu_1, \nu_2)}{\partial \underline{x}_1} \cdot \left(\frac{\partial f_2}{\partial \underline{p}_1} - \frac{\partial f_2}{\partial \underline{p}_2} \right)} - \boxed{\frac{\partial \mathcal{W}(|\underline{x}_1 - \underline{x}_2|, \nu_1, \nu_2)}{\partial \nu_1} \cdot \left(\frac{\partial f_2}{\partial \underline{\zeta}_1} - \frac{\partial f_2}{\partial \underline{\zeta}_2} \right)} = 0. \end{aligned}$$

BOLTZMANN–CURTISS EQUATION

4



J. Chem. Phys. 1956, 24, 225–241 (C. F. Curtiss),
J. Chem. Phys. 1963, 38, 2352–2363 (C. F. Curtiss, J. S. Dahler).

We can obtain from the embedded BBGKY hierarchy the following **Boltzmann** type equation,

$$\partial_t f + \nabla_{\mathbf{x}} \cdot (\underline{\mathbf{v}}f) + \nabla_{\underline{\alpha}} \cdot (\underline{\dot{\alpha}}f) = C[f, f] \quad (1)$$

where $f(\mathbf{x}, \underline{\mathbf{v}}, \underline{\alpha}, \underline{\zeta})$ is the probability of having a particle at $(\mathbf{x}, \underline{\mathbf{v}}, \underline{\alpha}, \underline{\zeta})$ in configuration space, normalised by $\frac{1}{n}$.



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We can obtain from the embedded BBGKY hierarchy the following **Boltzmann** type equation,

$$\partial_t f + \nabla_{\mathbf{x}} \cdot (\underline{v}f) + \nabla_{\underline{\alpha}} \cdot (\dot{\underline{\alpha}}f) = C[f, f]$$

where $f(\mathbf{x}, \underline{v}, \underline{\alpha}, \underline{s})$ is the probability of having a particle at $(\mathbf{x}, \underline{v}, \underline{\alpha}, \underline{s})$ in configuration space, normalised by $\frac{1}{n}$.

$$C[f, f] = \iiint\int (f'_1 f' - f_1 f)(\underline{k} \cdot \underline{g}) S(\underline{k}) d\underline{k} d\underline{v}_1 d\underline{\alpha}_1 d\underline{s}_1$$

with $S(\underline{k}) d\underline{k}$ being the surface element of the excluded volume and $\underline{g} = \underline{v}_1 - \underline{v} + \underline{\omega}_1 \times \mathbf{x}_1 - \underline{\omega} \times \mathbf{x}$.

COLLISION INVARIANTS



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It is possible to prove that the following quantities are **collision invariants** for $C[f, f]$, i.e.

$$\iiint \psi^{(i)} C[f, f] d\underline{v}_1 d\underline{\zeta}_1 d\underline{\alpha}_1 = 0.$$

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- ▶ $\psi^{(4)} = \frac{1}{2}m\underline{v} \cdot \underline{v} + \frac{1}{2}\underline{\omega} \cdot \mathbb{I} \cdot \underline{\omega}$, the **kinetic energy of the system**.

THE HYDRODYNAMIC EQUATIONS – NOTATION



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We first introduce the **number density**, i.e.

$$n(\mathbf{x}) = \iiint f(\mathbf{x}, \underline{v}, \underline{\alpha}, \underline{\zeta}) d\underline{v} d\underline{\alpha} d\underline{\zeta}.$$

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Then we can give a meaning to the following *chevrons*, i.e.

$$\langle\langle \cdot \rangle\rangle(\mathbf{x}) := \frac{1}{n(\mathbf{x})} \iiint \cdot f(\mathbf{x}, \underline{v}, \underline{\alpha}, \underline{\zeta}) d\underline{v} d\underline{\alpha} d\underline{\zeta}.$$

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Using this notation we can define **macroscopic stream velocity** and **macroscopic stream angular velocity** respectively as:

$$\underline{v}_0 := \langle\langle \underline{v} \rangle\rangle, \quad \underline{\omega}_0 := \langle\langle \underline{\omega} \rangle\rangle.$$

THE HYDRODYNAMIC EQUATIONS – CURTISS BALANCE LAWS



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Testing (1) against the first two **collision invariants** and integrating, Curtiss obtained the following **balance laws**:

$$\begin{aligned}\partial_t \rho + \nabla_{\mathbf{x}} \cdot (\rho \underline{v}_0) &= 0, \\ \rho \left[\partial_t \underline{v}_0 + (\nabla_{\mathbf{x}} \underline{v}_0) \underline{v}_0 \right] + \nabla_{\mathbf{x}} \cdot (\rho \mathbb{P}) &= 0,\end{aligned}$$

where ρ is the **density** defined as $\rho(\mathbf{x}) = mn(\mathbf{x})$ and \mathbb{P} is the **pressure tensor** defined as $\mathbb{P} := \langle\langle \underline{V} \otimes \underline{V} \rangle\rangle$, with \underline{V} being the **peculiar velocity** $\underline{V} := \underline{v} - \underline{v}_0$.

THE HYDRODYNAMIC EQUATIONS – SURPRISE BALANCE LAWS



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For the third collision invariant we took a different route than Curtiss, which led to the following balance law

$$\rho \left[\partial_t \underline{\eta} + (\nabla_{\mathbf{x}} \underline{\eta}) \underline{v}_0 \right] + \nabla_{\mathbf{x}} \cdot (\rho \mathbb{N}) = \underline{\xi},$$

where $\underline{\eta}$ is the **macroscopic intrinsic angular momentum** defined as $\underline{\eta}(\mathbf{x}) := \langle\langle \mathbb{I} \cdot \omega \rangle\rangle$ and \mathbb{N} is the **couple tensor** defined as $\mathbb{N} := \langle\langle \underline{V} \otimes (\mathbb{I} \underline{\omega}) \rangle\rangle$. Here ξ_l is defined in tensor notation as $\langle\langle mn(\varepsilon_{lki} v_i v_k) \underline{e}_l \rangle\rangle$ and we proved that $\underline{\xi}$ vanishes.

MAXWELL–BOLTZMANN DISTRIBUTION



J. Chem. Phys. 1956, 24, 225–241 (C. F. Curtiss),
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Curtiss gives an expression for the Maxwell–Boltzmann distribution, i.e. the distribution $f^{(0)}$ such that $C[f^{(0)}, f^{(0)}]$ vanishes:

$$f^{(0)}(\underline{v}, \underline{\omega}) = \frac{n \sin(\alpha_2) Q}{\int Q \sin(\alpha_2) d\underline{\alpha}} \frac{m^{\frac{3}{2}}}{(2\pi \langle\langle \theta \rangle\rangle)^3} (\Gamma)^{\frac{1}{2}} \exp \left[-m \frac{|\underline{V}|}{2 \langle\langle \theta \rangle\rangle} - \frac{\underline{\Omega} \cdot \mathbb{I} \cdot \underline{\Omega}}{2 \langle\langle \theta \rangle\rangle} \right],$$

where the **peculiar angular velocity** defined as $\underline{\Omega} := \underline{\omega} - \underline{\omega}_0$, $\Gamma = \prod_{i=1}^3 \Gamma_i$, Γ_i are the moments of inertia of the spherocylinder we are considering and $Q := \exp \left[\frac{\underline{\omega}_0 \cdot \mathbb{I} \cdot \underline{\omega}_0}{2\theta} \right]$.

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Notice in particular that we assumed $\underline{\omega}_0$ is the kinetic temperature of the system measured in energy units.

MOMENTUM CLOSURE AROUND THE EQUILIBRIUM

Now we can use the Maxwell–Boltzmann distribution to compute an approximation of the **pressure tensor** near the equilibrium, i.e.

$$\mathbb{P}^{(0)} = \frac{\Gamma}{3m} \langle\langle \theta \rangle\rangle \underline{Id}.$$

$$\left[\partial_t \underline{v}_0 + (\nabla_{\mathbf{x}} \underline{v}_0) \underline{v}_0 \right] = -\frac{1}{\rho} \nabla p,$$

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Unfortunately the same procedure results in a **vanishing** $\mathbb{N}^{(0)}$.

NOETHER'S THEOREM AND MOMENTUM COUPLING

Let us consider the equation for the angular momentum, and observe that under the assumption $\mathbb{N}^{(0)} = 0$ it reads

$$\underline{\dot{\eta}} = \underline{\xi} = 0.$$

In particular, this is a consequence of Noether's theorem since when $\mathbb{N}^{(0)} = 0$ we have a **rotationally invariant** Lagrangian.

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Near the thermal equilibrium is the fluid isotropic? No!

THE NEMATIC ORDERING AND THE INERTIA TENSOR

We know that for a **slender body** the inertia tensor can be decomposed as,

$$\mathbb{I} = \lambda_1(I - \underline{\nu} \otimes \underline{\nu}) + \mathcal{O}(\varepsilon)$$

where $\varepsilon = (\frac{r}{a})^2$. Furthermore, the macroscopic kinetic energy can be computed as

$$m \frac{1}{2} |\underline{v}_0|^2 + \frac{1}{2} \underline{\omega}_0 \cdot \mathbb{I} \underline{\omega}_0 = \frac{1}{2} m |\underline{v}_0|^2 + \frac{\lambda_1}{2} |\dot{\underline{\nu}}|^2 + \mathcal{O}(\varepsilon),$$

as $\varepsilon \rightarrow 0$ we retrieve the same energy that is the starting point for **Ericksen theory of anisotropic fluids**.

BALANCE LAWS FOR KINETIC TEMPERATURE



Multiscale Model. Simul. 2024, accepted (P. E. Farrell, G. Russo,),

We need another way to formulate the **constitutive relation** for the **couple tensor**. We begin by observing that from $\psi^{(4)}$ we get the following balance law:

$$\dot{\psi}_0 + \nabla_{\mathbf{x}} \underline{v}_0 : (\rho \mathbb{P}) + \nabla_{\mathbf{x}} \underline{\omega}_0 : (\rho \mathbb{N}) - \nabla \cdot \left[\mathbb{P}^T \underline{v}_0 + \mathbb{N}^T \underline{\omega}_0 \right] + \nabla_{\mathbf{x}} \cdot \mathbf{Q} = 0$$

where $\psi_0 = \langle\langle \theta \rangle\rangle$, $\mathbf{Q} = \frac{1}{2} \langle\langle \underline{V}(m|\underline{V}|^2 + \underline{\Omega} \cdot \mathbb{I} \underline{\Omega}) \rangle\rangle$, and

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$$\theta = \frac{m}{2} \underline{V} \cdot \underline{V} + \frac{1}{2} \underline{\Omega} \cdot \mathbb{I} \cdot \underline{\Omega}.$$

This is a **kinetic derivation** of **Leslie's rate of work hypothesis**.

THE OSEEN-FRANK STORED ENERGY


Multiscale Model. Simul. 2024, accepted (P. E. Farrell, G. Russo, U. Z.),

Making use of the fact that $\dot{\underline{\nu}} = \underline{\omega} \times \underline{\nu} = \partial_t \underline{\nu}(\nabla \underline{\nu}) \underline{\nu}$ we can rewrite part of the stored energy as

$$\psi_{OF}(\underline{\nu}, \nabla \underline{\nu}) = \frac{1}{2} \underline{\Omega} \cdot \mathbb{I} \underline{\Omega} = \frac{\lambda_1}{2} \text{tr} \left[\nabla \underline{\nu}^T \mathbb{P}^{(0)} \nabla \underline{\nu} \right].$$

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Using $\mathbb{P}^{(0)}$ we get a **stored energy functional** very similar to the **Oseen-Frank** energy

$$\psi_{OF} = p \frac{\lambda_1}{2} \text{tr} \left[\nabla \underline{\nu}^T \nabla \underline{\nu} \right].$$

NOLL–COLEMAN PROCEDURE



Multiscale Model. Simul. 2024, accepted (P. E. Farrell, G. Russo, U. Z.),

Since we are happy with our **pressure tensor**, we make the following **ansatz**

$$\psi = \psi(\nu, \nabla \nu)$$

where ν is the **nematic director**. Expanding the total derivative and using the Ericksen identity we get the following expression in tensor notation

$$\dot{\psi} = \varepsilon_{iqp} \left[(\nu_q \frac{\partial \psi}{\partial (\nu_p)} + \partial_k (\nu_q) \frac{\partial \psi}{\partial (\partial_k \nu_p)}) \omega_i^0 + \nu_q \frac{\partial \psi}{\partial (\partial_k \nu_p)} \partial_k \omega_i^0 \right] - \frac{\partial \psi}{\partial (\partial_k \nu_p)} \partial_q (\nu_p) \partial (v_q^0)$$

NOLL–COLEMAN PROCEDURE



Multiscale Model. Simul. 2024, accepted (P. E. Farrell, G. Russo, U. Z.),

Substituting this expression in the Theorem of Power Expended and considering thermodynamic processes for which the exact divergences disappear, we get:

$$\left[\mathbb{P}_{ij} + \frac{\partial \psi}{\partial (\partial_j \nu_p)} \partial_i (\nu_p) \right] \partial_j (\nu_i) + \left[N_{ij} - \varepsilon_{iqp} \nu_q \frac{\partial \psi}{\partial (\partial_j \nu_p)} \right] \partial_j (\omega_i^0)$$

$$\left[P_{pq} - \frac{\partial \psi}{\partial (\partial_p \nu_k) \partial_q (\nu_k)} \right] \varepsilon_{iqp} \omega_i^0 \geq 0.$$

Since the above expression must hold for all thermodynamic processes for which the exact divergences disappear, we get the following **constitutive relations**:

$$\mathbb{P} = \nabla \underline{\nu}^T \frac{\partial \psi}{\partial (\nabla \underline{\nu})} + \mathbb{P}^{(0)}, \quad \mathbb{N}_{ij} = \varepsilon_{iqp} \nu_q \frac{\partial \psi}{\partial (\partial_j \nu_p)} = \underline{\nu} \times \frac{\partial \psi}{\partial (\nabla \underline{\nu})}.$$

COMPRESSIBLE LESLIE–ERICKSEN EQUATIONS



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This leads to the following set of equations, which can be seen as an inviscid compressible generalisation of the Leslie–Ericksen equations:

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VLASOV–TYPE EQUATION

arXiv 2025, (J. A. Carillo, P. E. Farrell, A. Medaglia,).

From the separation of timescales in the BBGKY hierarchy we obtain the following identity,

$$\frac{\underline{p}_2 - \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \mathbf{x}} = \frac{\partial \mathcal{W}}{\partial \mathbf{x}_1} (|\mathbf{x}_1 - \mathbf{x}_2|, \nu_1, \nu_2) \cdot \left(\frac{\partial f_2}{\partial \underline{p}_1} - \frac{\partial f_2}{\partial \underline{p}_2} \right).$$

Substituting this identity in the second equation of the BBGKY hierarchy we obtain the following equation,

$$\begin{aligned} \frac{\partial f_1}{\partial t} + \frac{\underline{p}_1}{m} \cdot \frac{\partial f_1}{\partial \mathbf{x}_1} + \underline{\underline{\Omega}}(\nu_1)^{-1} \underline{\underline{\zeta}}_1 \cdot \frac{\partial f_1}{\partial \nu_1} &= \int \frac{\underline{p}_2 - \underline{p}_1}{m} \cdot \frac{\partial f_2}{\partial \mathbf{x}} d\Gamma_2 \\ &+ \int \frac{\partial \mathcal{W}(|\mathbf{x}_1 - \mathbf{x}_2|, \nu_1, \nu_2)}{\partial \nu_1} \cdot \frac{\partial f_2}{\partial \underline{\zeta}_1} d\Gamma_2. \end{aligned}$$

WEAK-ORDER INTERACTIONS

We might be tempted to assume interactions are **weak**,

$$f_2(\Gamma_1, \Gamma_2, t) = f_1(\Gamma_1, t)f_1(\Gamma_2, t).$$

This leads to equations of a **reversible nature**, compatible with **Loschmidt's paradox**.

Thus, we have no guarantee that the system described thermalises to a Maxwellian distribution.

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Thus, we have no guarantee that the system described thermalises to a Maxwellian distribution.

Weak-order Interactions

We will say that a kinetic equation is governed by **weak-order interactions** if the derivative of the two-particle distribution function factorises as,

$$\partial_{v_i} f_2(\Gamma_1, \Gamma_2, t) = f_1(\Gamma_i, t) \partial_{v_i} f_1(\Gamma_j, t),$$

$$\partial_{c_i} f_2(\Gamma_1, \Gamma_2, t) = f_1(\Gamma_j, t) \partial_{c_i} f_1(\Gamma_i, t),$$

for $i \neq j$ and $i, j = 1, 2$.

VLASOV–TYPE EQUATION

Under the assumption of weak-order interactions we can rewrite the first equation of the BBGKY hierarchy as,

$$\frac{\partial f}{\partial t} + \dot{\underline{x}} \cdot \nabla_{\underline{x}} f + \dot{\underline{\nu}} \cdot \nabla_{\underline{\nu}} f + \mathcal{V} \cdot \nabla_{\underline{\varsigma}} f = C[f, f],$$

where the collision operator $C[f, f]$ can be written using the transition “probability” W as,

$$C[f_1, f_1] = \int d\underline{\Xi}'_1 d\underline{\Xi}'_2 d\underline{\Xi}_2 \int_0^{\frac{\pi}{2}} \int_0^{2\pi} W(\underline{\Xi}'_1, \underline{\Xi}'_2 \mapsto \underline{\Xi}_1, \underline{\Xi}_2) f_1(\Gamma'_1, t) f_1(\Gamma'_2, t) \\ - W(\underline{\Xi}_1, \underline{\Xi}_2 \mapsto \underline{\Xi}'_1, \underline{\Xi}'_2) f_1(\Gamma_1, t) f_1(\Gamma_2, t) d\theta_2 d\varphi_2. \\ \mathcal{V}(\underline{\mathbf{x}}_1, \nu_1, t) = \iiint \frac{\partial \mathcal{W}(\underline{\mathbf{x}}_1 - \underline{\mathbf{x}}_2, \nu_1, \nu_2)}{\partial \nu_1} f(\underline{\mathbf{x}}_2, \nu_2, \varsigma_2, t) d\underline{\mathbf{x}}_2 d\nu_2 d\varsigma_2.$$

BOLTZMANN INEQUALITY AND THERMALISATION



J. Stat. Phys. Volume 26, 795–801 (C. Cercignani, M. Lampis),
arXiv 2025, (J. A. Carillo, P. E. Farrell, A. Medaglia,).

As we said before the collision operator $C[f, f]$ considered here guarantees that the system thermalises to a Maxwellian distribution. In particular, we can prove

$$\int d\Xi \log(f(\Gamma, t)) C[f, f] \leq 0,$$

which is a generalisation of the **Boltzmann inequality** for Boltzmann's equation with internal degrees of freedom. Following the classical calculus of variation approach we can prove that the unique Maxwellian with prescribed collision invariants is

$$\bar{f}(\Gamma, t) = \exp \left(a + \underline{b} \cdot \underline{p} + c(\underline{p} \times \underline{x} + \underline{w}_\nu \times \underline{s}) + d(m^{-1} \underline{p} \cdot \underline{p} + \varsigma \cdot \underline{\underline{\Omega}}(\nu)^{-1} \varsigma) \right).$$

SPACE HOMOGENEOUS VLASOV-TYPE EQUATION

We are interested in the time evolution of the distribution $f(\underline{v}, \nu, \underline{\zeta}, t)$, $\underline{v} \in \mathbb{R}^2$, $\nu \in \mathcal{M}$, $\underline{\zeta} \in T_\nu \mathcal{M}$, and $t \geq 0$, solution to the space-homogeneous equation

$$\frac{\partial f}{\partial t} + \underline{\underline{\Omega}}(\nu)^{-1} \underline{\zeta} \cdot \nabla_\nu f + \mathcal{V} \cdot \nabla_{\underline{\zeta}} f = \frac{1}{\tau} \mathcal{C}[f, f],$$

where τ has been obtained rescaling the collision frequency, and as collision operator we consider the one associated with Maxwellian molecules, i.e.

$$\mathcal{C}[f, f] = \int d\underline{\zeta}_2 d\underline{v}_2 d\nu_2 f' f'_* - \int d\underline{\zeta}_2 d\underline{v}_2 d\nu_2 f f_*, \quad (1)$$

complemented with initial conditions $f(\underline{v}, \nu, \underline{\zeta}, 0) = f_0(\underline{v}, \nu, \underline{\zeta})$ and where we will denote $f_* = f(\underline{v}_2, \nu_2, \underline{\zeta}_2, t)$, and f', f'_* are the distributions depending on the post interaction coordinates.

DIRECT SIMULATIONS MONTE CARLO (DSMC)

We consider discretization of the time interval $[0, T_f]$, with $T_f > 0$ final simulation time, of step $\Delta t > 0$ such that $t^n = n\Delta t$. By $f^n(\underline{v}, \nu, \underline{\zeta})$ we denote an approximation of $f(\underline{v}, \nu, \underline{\zeta}, t^n)$ at the n -th time step and we apply a splitting method between the Vlasov-type transport operator and the collisional operator.

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Transport $\mathcal{T}_{\Delta t}(\cdot)$

We solve the Vlasov-type step $\hat{f} = \mathcal{T}_{\Delta t}(f^n)$

$$\begin{cases} \frac{\partial \hat{f}}{\partial t} + \Omega(\nu)^{-1} \underline{\zeta} \cdot \nabla_{\nu} \hat{f} + \mathcal{V} \cdot \nabla_{\underline{\zeta}} \hat{f} = 0 \\ \hat{f}(\underline{v}, \nu, \underline{\zeta}, 0) = f^n(\underline{v}, \nu, \underline{\zeta}) \end{cases}$$

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Collision $\mathcal{Q}_{\Delta t}(\cdot)$

We then solve the collision step $\hat{\hat{f}} = \mathcal{Q}_{\Delta t}(\hat{f})$ with initial data given by the solution of the previous step

$$\begin{cases} \tau \frac{\partial \hat{\hat{f}}}{\partial t} = \mathcal{C}[\hat{\hat{f}}, \hat{\hat{f}}] \\ \hat{\hat{f}}(\underline{v}, \nu, \underline{\zeta}, 0) = \hat{f}(\underline{v}, \nu, \underline{\zeta}, \Delta t). \end{cases}$$

DIRECT SIMULATIONS MONTE CARLO (DSMC)

We consider discretization of the time interval $[0, T_f]$, with $T_f > 0$ final simulation time, of step $\Delta t > 0$ such that $t^n = n\Delta t$. By $f^n(\underline{v}, \nu, \underline{\zeta})$ we denote an approximation of $f(\underline{v}, \nu, \underline{\zeta}, t^n)$ at the n -th time step and we apply a splitting method between the Vlasov-type transport operator and the collisional operator.

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We solve the Vlasov-type step $\hat{f} = \mathcal{T}_{\Delta t}(f^n)$

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Collision $\mathcal{Q}_{\Delta t}(\cdot)$

We then solve the collision step $\hat{\hat{f}} = \mathcal{Q}_{\Delta t}(\hat{f})$ with initial data given by the solution of the previous step

$$\begin{cases} \tau \frac{\partial \hat{\hat{f}}}{\partial t} = \mathcal{C}[\hat{\hat{f}}, \hat{\hat{f}}] \\ \hat{\hat{f}}(\underline{v}, \nu, \underline{\zeta}, 0) = \hat{f}(\underline{v}, \nu, \underline{\zeta}, \Delta t). \end{cases}$$

The first order in time splitting finally reads $f^{n+1}(\underline{v}, \nu, \underline{\zeta}) = \mathcal{Q}_{\Delta t}(\mathcal{T}_{\Delta t}(f^n)(\underline{v}, \nu, \underline{\zeta}))$.

DSMC: TRANSPORT $\mathcal{T}_{\Delta t}(\cdot)$

We introduce an approximation of the distribution function with a sample of N particles identified by their velocities \underline{v}_i^n , order parameter ν_i^n , and conjugate momentum $\underline{\varsigma}_i^n$ at the time t^n , for $i = 1, 2, \dots, N$,

$$f^n(\underline{v}, \nu, \underline{\varsigma}) \approx f^{n,N}(\nu, \underline{\varsigma}) = \sum_{i=1}^N \delta(\nu - \nu_i(t^n)) \otimes \delta(\underline{\varsigma} - \underline{\varsigma}_i(t^n)).$$

The Vlasov-type transport step $\mathcal{T}_{\Delta t}(\cdot)$ is solved by considering the characteristic equations associated to the operator, which as discussed in the previous section, result in a system of (time-continuous) ODEs

$$\frac{d\underline{v}_i}{dt} = \underline{\varsigma}_i, \quad \frac{d\underline{\varsigma}_i}{dt} = \mathcal{V}(\nu_i, \underline{\varsigma}_i).$$

This system is solved, at the time discrete level, with a classical first order semi-implicit Euler scheme for the time derivative.

DSMC: COLLISION $\mathcal{Q}_{\Delta t}(\cdot)$

The collisional step $\mathcal{Q}_{\Delta t}(\cdot)$ is solved with a classical Nanbu-Babovsky DSMC approach. First, we rewrite the collisional operator to highlight the gain and loss part integrating the second term in (1)

$$G - L = \int d\underline{s}_2 d\underline{v}_2 d\nu_2 f' f'_* - f,$$

and then we discretize the time derivative with a first order in time Euler scheme to obtain

$$f^{n+1} = \left(1 - \frac{\Delta t}{\tau}\right) f^n + \frac{\Delta t}{\tau} \int d\underline{s}_2 d\underline{v}_2 d\nu_2 f' f'_*.$$

We have thus rewritten f^{n+1} as a convex combination of f^n and the gain term, i.e. we will consider all the particles in the system with probability $\frac{\Delta t}{\tau}$ we will update the velocity, order parameter and conjugate momentum according to the binary law relating the pre and post interaction velocities, order parameters and conjugate momenta.

AN NUMERICAL EXAMPLE: ROD-LIKE MOLECULES

In the context of rod-like molecules, with vanishing girth, we can explicitly compute the Vlasov-type force \mathcal{V} and the transport term to obtain the following equation

$$\frac{\partial f}{\partial t} + \omega \nabla_{\theta} f + \mathcal{V} \cdot \nabla_{\omega} f = \iiint (f' f'_* - f f_*) dv_* d\theta_* d\omega_*,$$

where $f = f(\underline{v}, \theta, \omega, t)$, $f_* = f(\underline{v}_*, \theta_*, \omega_*, t)$, and f', f'_* are the distributions depending on the post interaction coordinates given by

$$\begin{aligned} \underline{v}' &= \underline{v} - (1 + e_v) \frac{J}{m} \underline{n}, & \underline{v}'_* &= \underline{v}_* + (1 + e_v) \frac{J}{m} \underline{n}, \\ \omega' &= \omega - (1 + e_{\omega}) J \mathbb{I}^{-1}(\underline{r} \times \underline{n}), & \omega'_* &= \omega_* + (1 + e_{\omega}) J \mathbb{I}_*^{-1}(\underline{r}_* \times \underline{n}), \end{aligned}$$

with

$$J = - \frac{\underline{V} \cdot \underline{n}}{\frac{2}{m} + [\mathbb{I}^{-1}(\underline{r} \times \underline{n}) \times \underline{r} + \mathbb{I}_*^{-1}(\underline{r}_* \times \underline{n}) \times \underline{r}_*] \cdot \underline{n}}.$$

Notice that $\theta' = \theta$ and $\theta'_* = \theta_*$ since the angles are not changed by the collisional operator.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

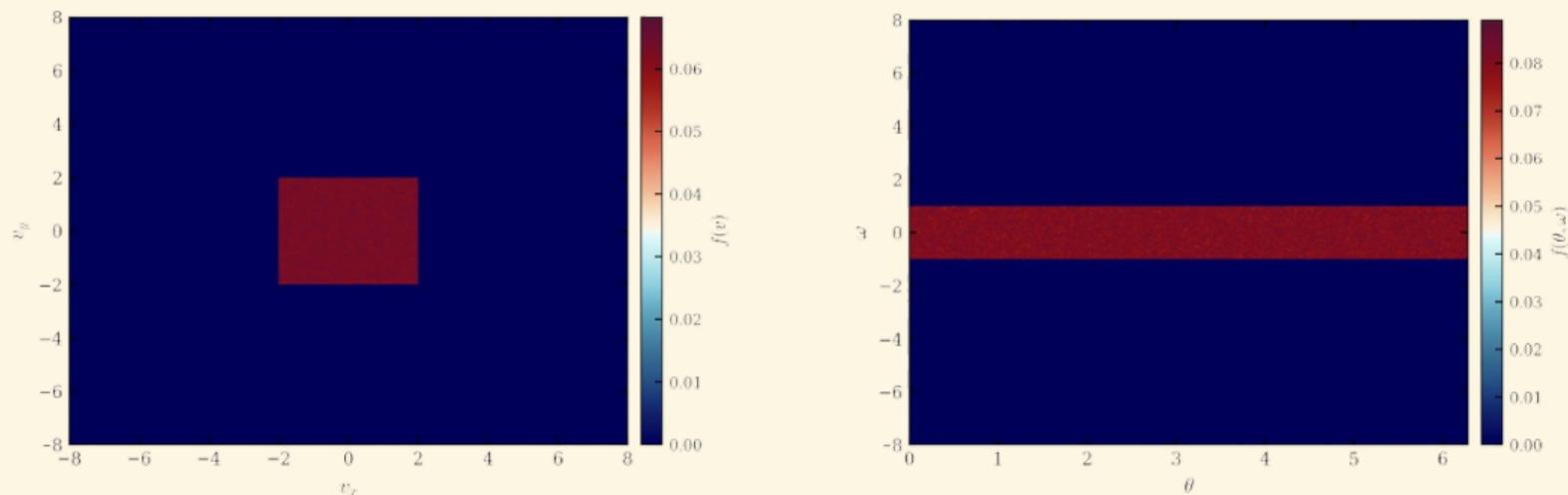


Figure: Test 1 - Zero Potential. The initial distribution function for the zero potential case. We consider a uniform initial velocity distribution (left) and a uniform distribution for the angular velocity with uniform distribution for the angle (right).

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

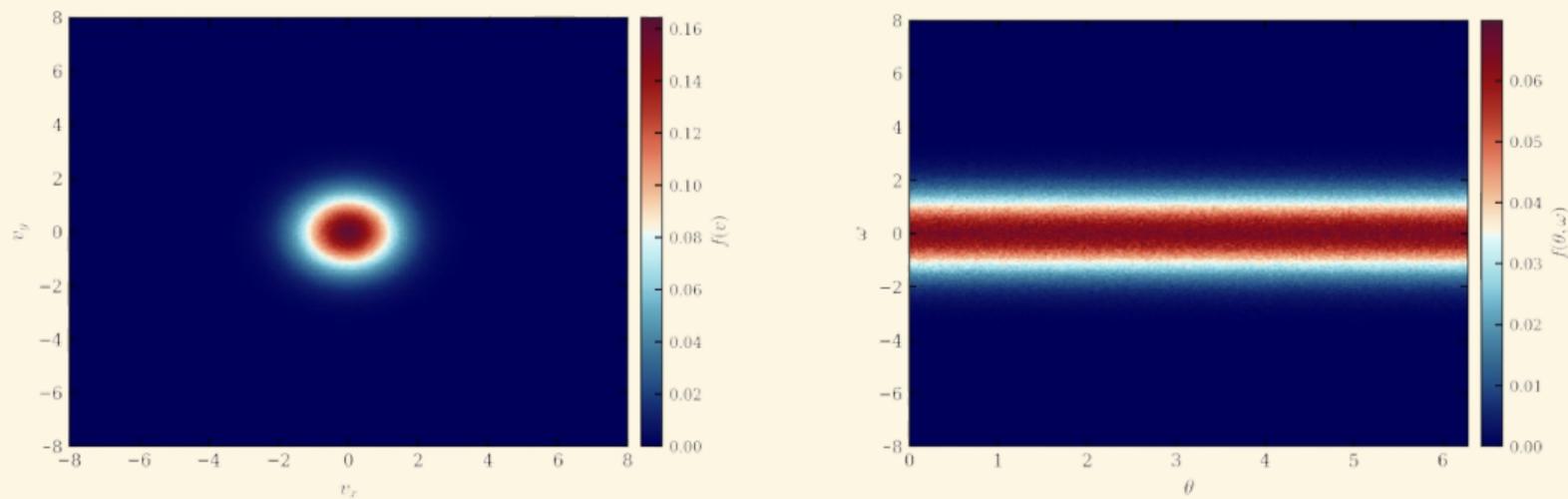


Figure: Test 1 - Zero Potential. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

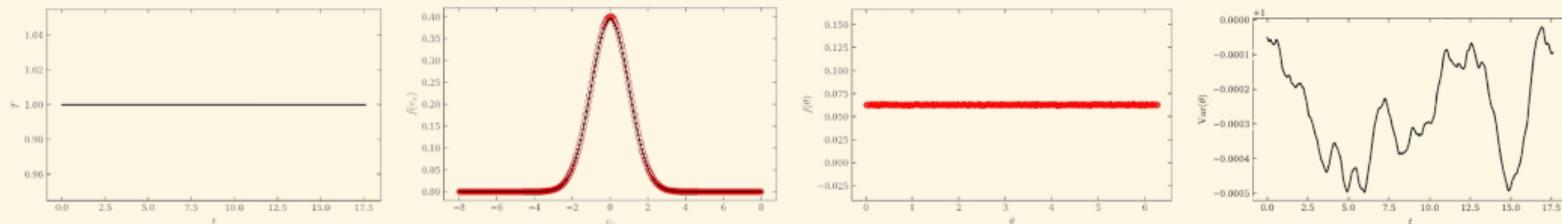


Figure: Test 1 - Zero Potential. The red circles the numerical results, the solid black lines the exact steady state. We observe a good accordance between numerical and exact results, the temperature (and energy) is conserved exactly, and the angle distribution remain a uniform, so we don't have aggregation.

Let us consider the mean-field potential is given by

$$\mathcal{W}(\underline{\nu}, \underline{\zeta}) = \frac{1}{2} \alpha (\underline{\nu} - \hat{\underline{\nu}}) \cdot (\underline{\nu} - \hat{\underline{\nu}}) + \beta \underline{\nu} \cdot \underline{\zeta}.$$

Under this hypothesis the Vlasov-type force can be computed to be

$$\mathcal{V}(\underline{\nu}, \underline{\zeta}) = -\alpha (\underline{\nu} - \hat{\underline{\nu}}) - \beta \underline{\zeta}.$$

This system of ODEs can be recasted as linear system of ODEs, i.e.

$$\begin{bmatrix} \frac{d\nu_i}{dt} \\ \frac{d\zeta_i}{dt} \end{bmatrix} = \begin{bmatrix} 0 & 1 \\ -\alpha & -\beta \end{bmatrix} \begin{bmatrix} \nu_i \\ \zeta_i \end{bmatrix} + \alpha \begin{bmatrix} 0 \\ \hat{\nu} \end{bmatrix}.$$

We can immediately see that the fixed points of the system is unique and it is given by $\underline{\nu} = \hat{\underline{\nu}}$ and $\underline{\zeta} = 0$. It remains to study the stability of the fixed point, which can be done by studying the eigenvalues of the Jacobian of the system which are given by

$$\lambda_{1,2} = \frac{-\beta \pm \sqrt{\beta^2 - 4\alpha}}{2}.$$

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

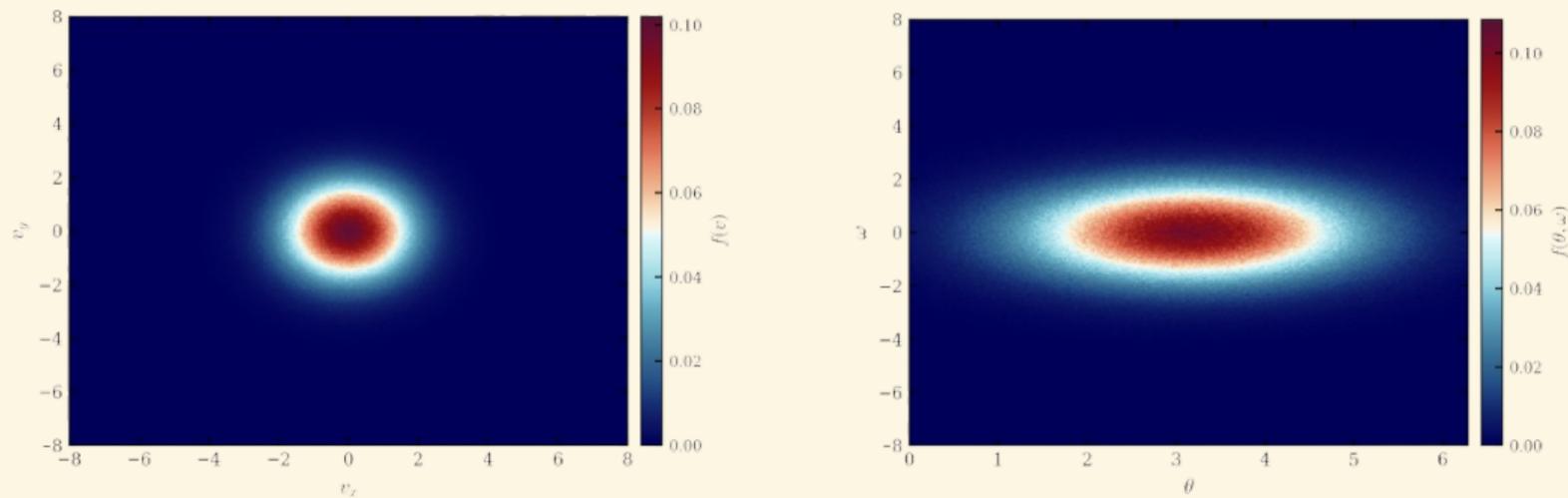


Figure: Test 2 - Quadratic potential for $\beta = 0$ and $\alpha = 1$. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

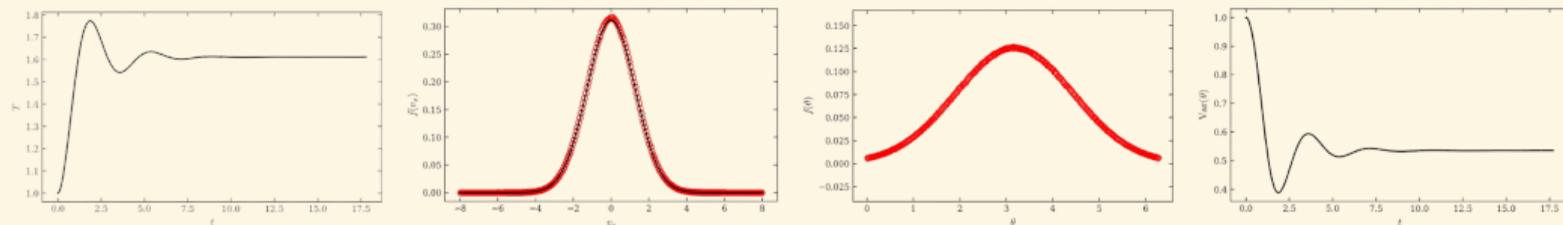


Figure: Test 2 - Quadratic potential for $\beta = 0$ and $\alpha = 1$. The red circles are the numerical results, the solid black lines the exact steady state. We observe a good accordance between numerical and exact results, and we do so the emergence of an alignment around the direction \hat{v} as expected from the stability analysis of the fixed point.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

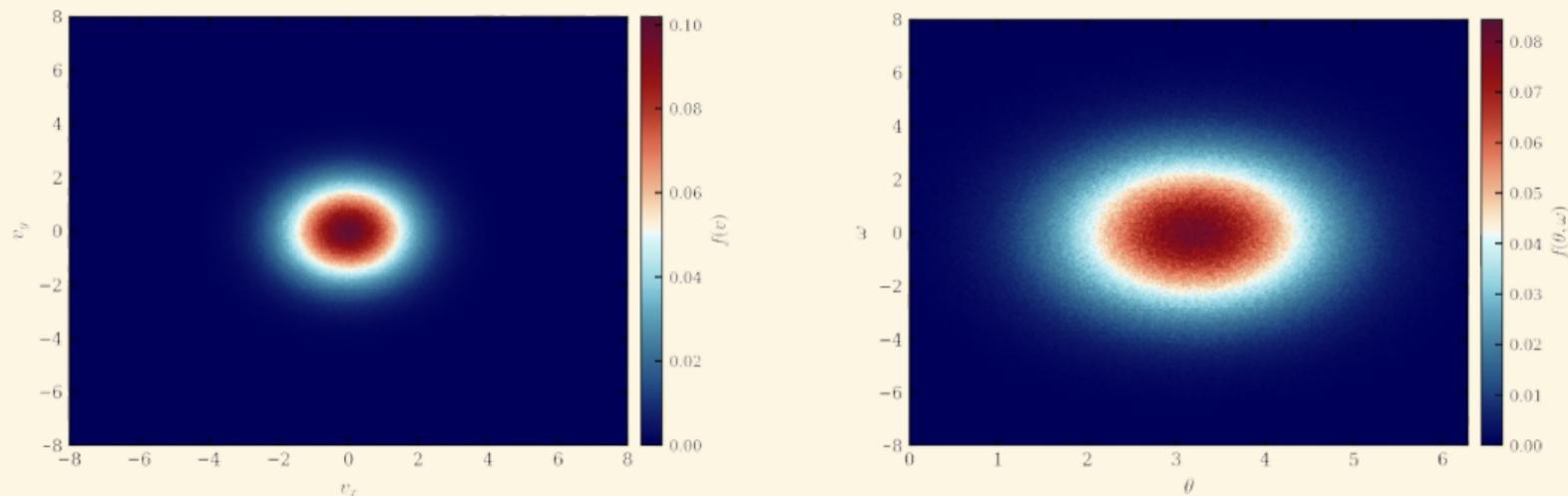


Figure: Test 3 - Quadratic potential for $\beta = 0$ and $\alpha = 4$. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

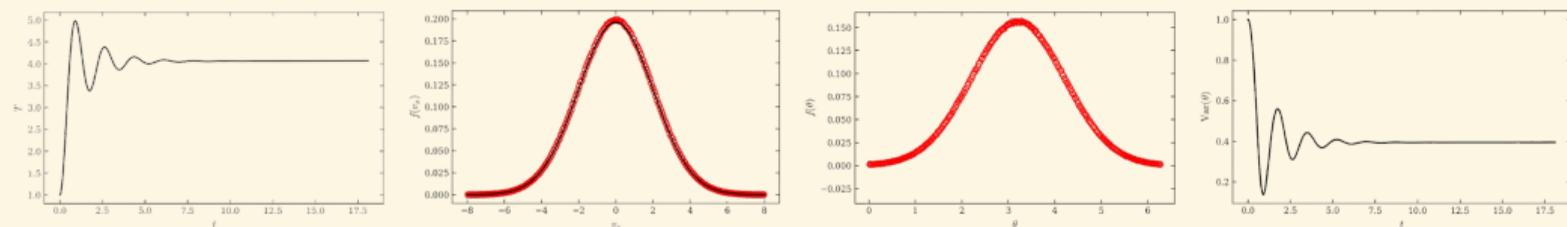


Figure: Test 3 - Quadratic potential for $\beta = 0$ and $\alpha = 4$. The red circles are the numerical results, the solid black lines the exact steady state. We observe a good accordance between numerical and exact results, and we do so the emergence of an alignment around the direction \hat{v} as expected from the stability analysis of the fixed point.

AN EXAMPLE: NEMATIC LIQUID CRYSTALS

We now consider a non-linear potential inspired by the moment of an electric dipole in an external field, i.e.

$$\mathcal{W}(\nu) = \alpha \sin(\theta - \hat{\theta}), \quad \theta = \arctan(\nu_x, \nu_y), \quad \hat{\theta} = \arctan(\hat{\nu}_x, \hat{\nu}_y),$$

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With this potential we see the alignment around random orientations resulting from the competition between the potential and the thermal agitation induced by the collisions.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

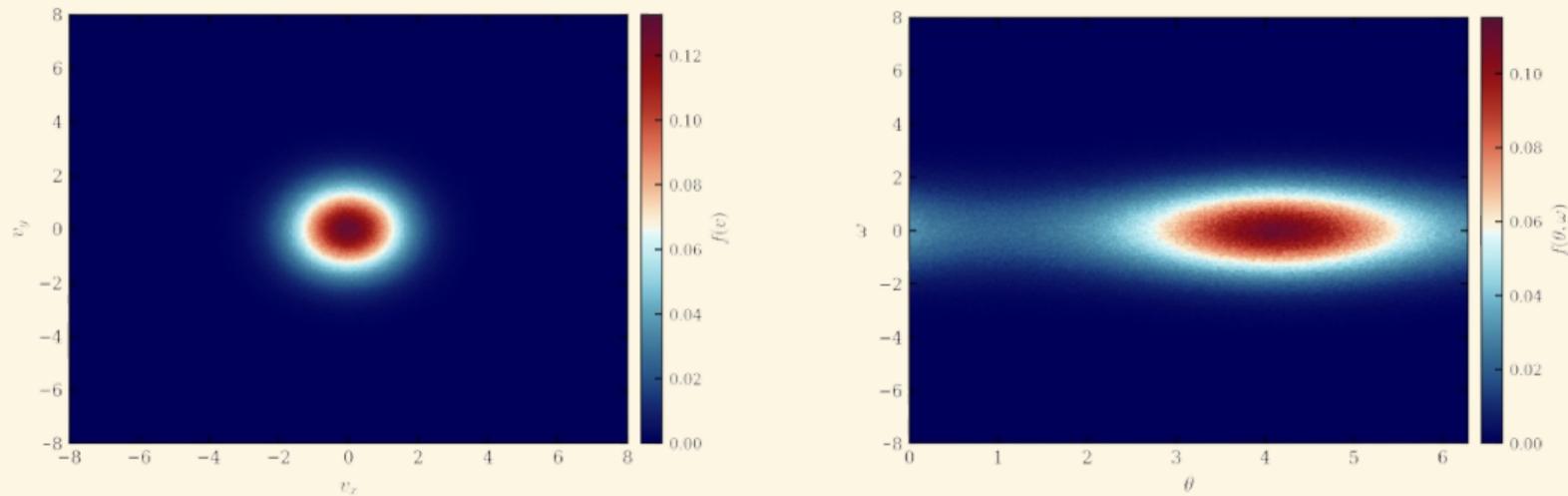


Figure: Test 4 – Non-linear potential for $\alpha = 1$, seed is 47. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

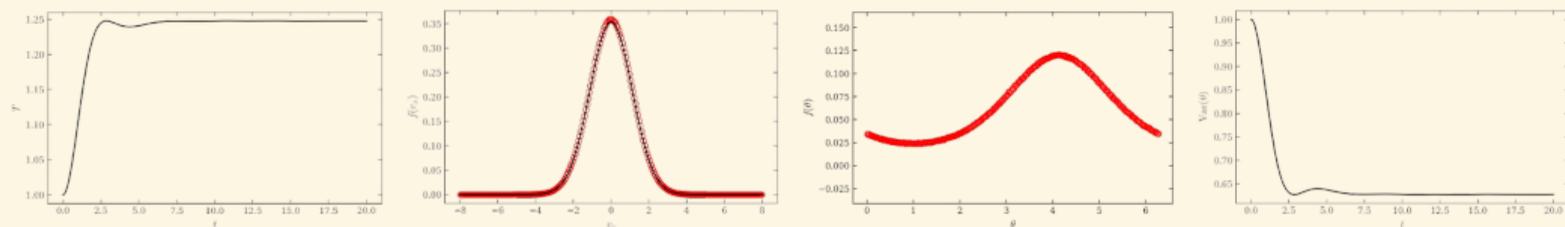


Figure: Test 4 – Non-linear potential for $\alpha = 1$, seed is 47. The red circles are the numerical results, the solid black lines the exact steady state. We observe a good accordance between numerical and exact results, and we do so the emergence of an alignment around the direction \hat{u} as expected from the stability analysis of the fixed point.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

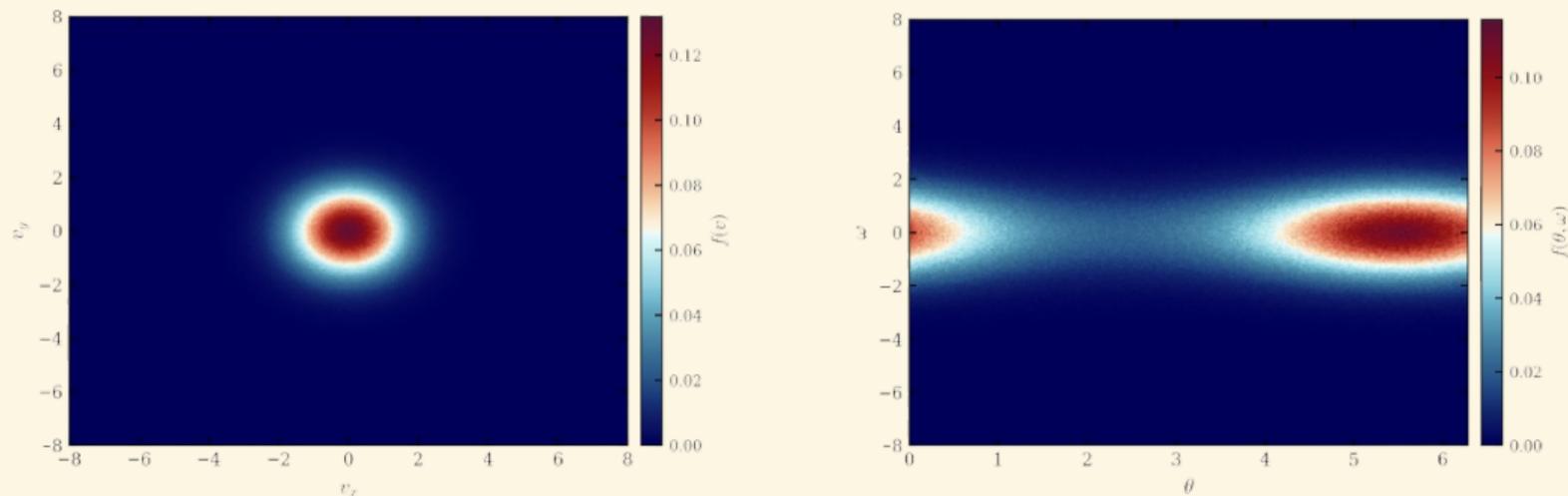


Figure: Test 6 – Non-linear potential for $\alpha = 1$, seed is 49. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

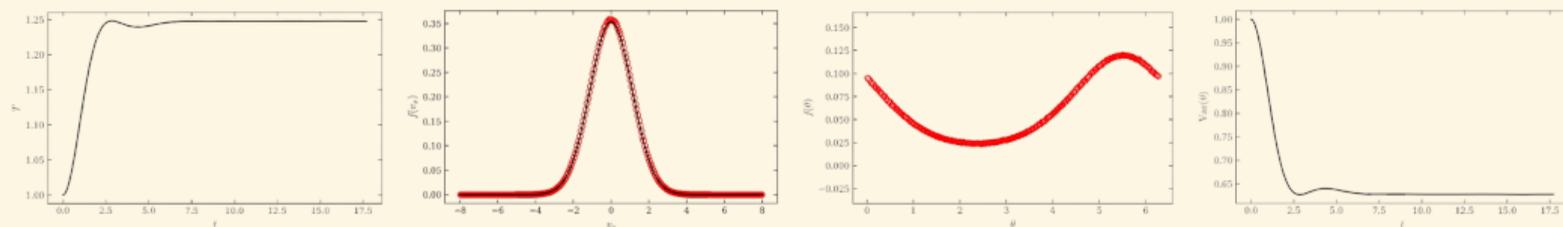


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AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

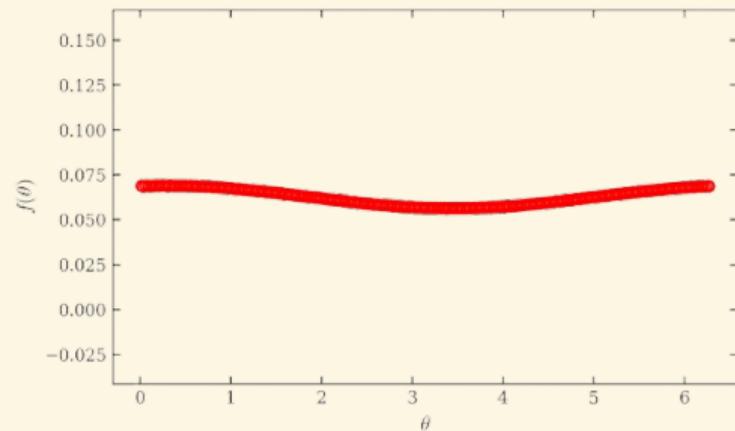
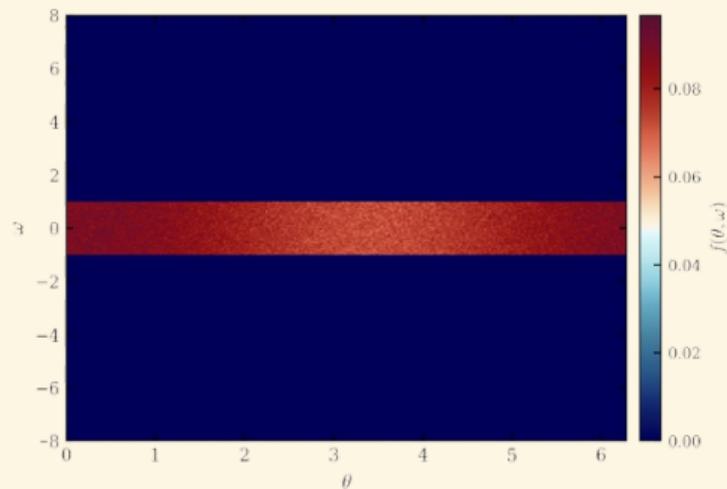


Figure: Test 5 – Non-linear potential for $\alpha = 1$, with nucleation. The initial distribution for the velocities and angular velocities remains unchanged with respect to the previous tests. The initial distribution for the angles is a perturbation of the uniform distribution, thus we can observe a nucleation process.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

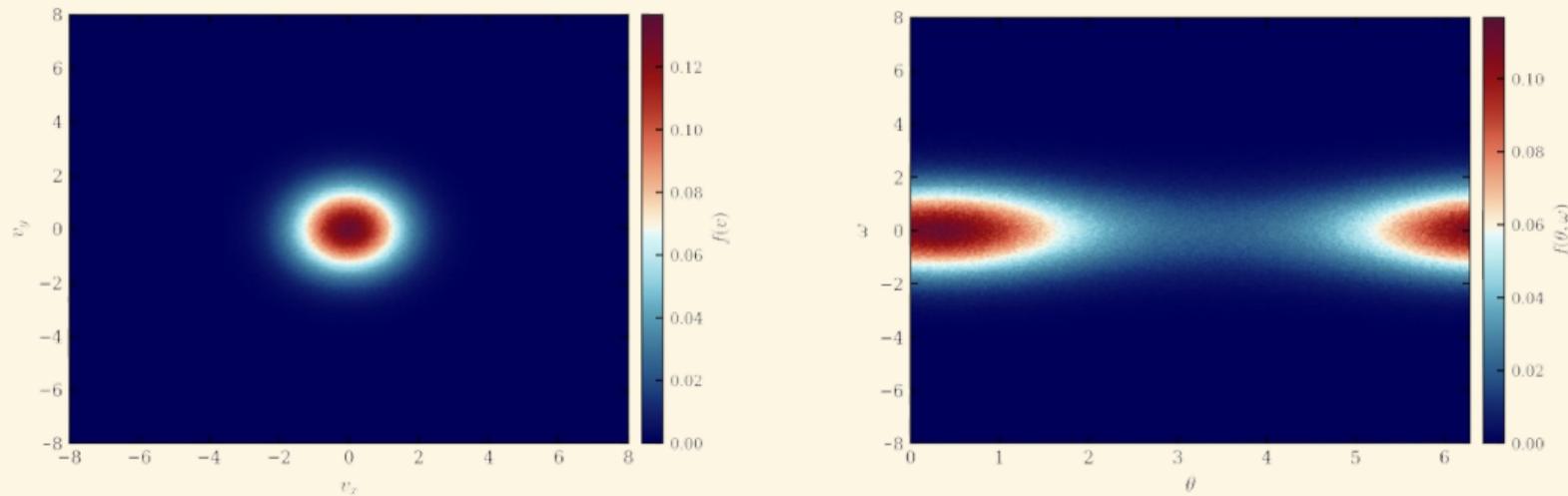


Figure: Test 5 – Non-linear potential for $\alpha = 1$, with nucleation. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

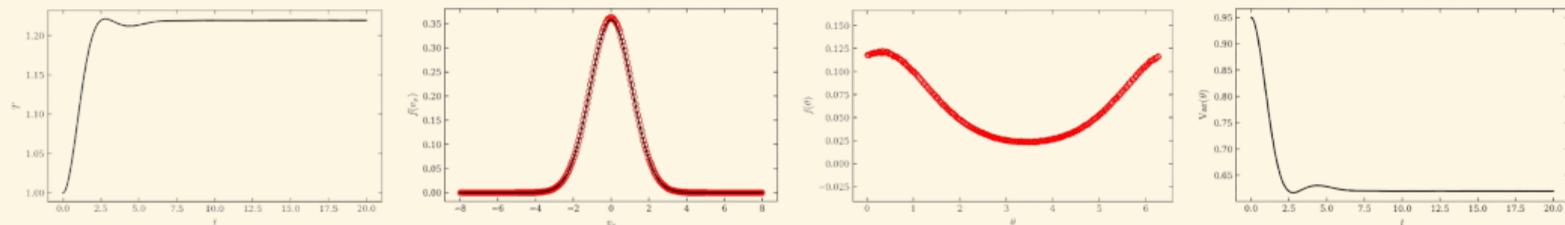


Figure: Test 5 – Non-linear potential for $\alpha = 1$, with nucleation. The red circles are the numerical results, the solid black lines the exact steady state. We observe a good accordance between numerical and exact results, and we do so the emergence of an alignment around the direction \hat{v} as expected from the stability analysis of the fixed point.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

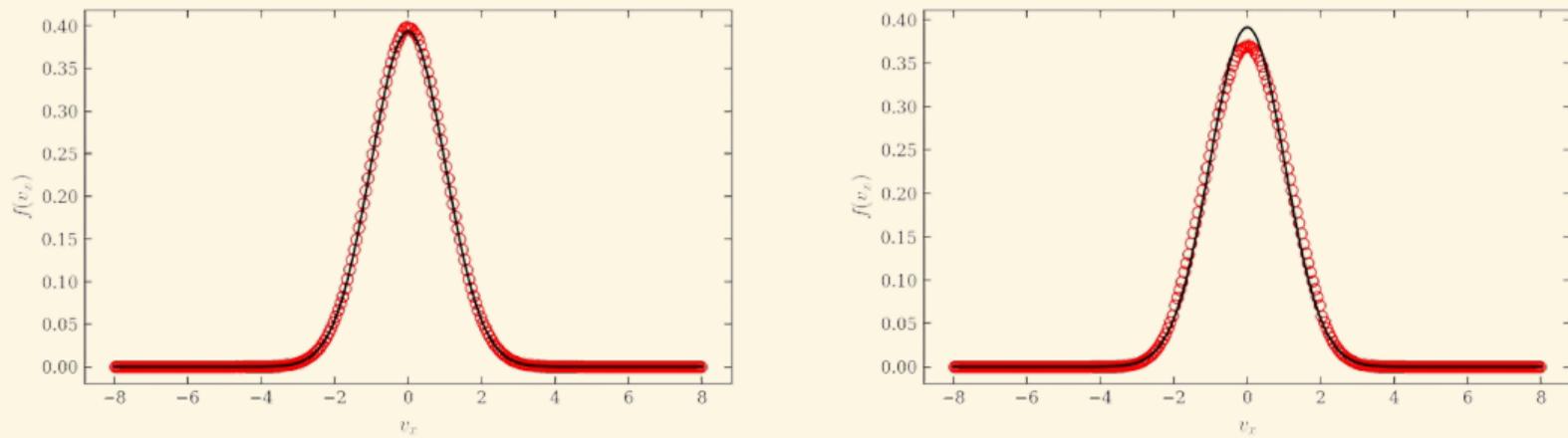


Figure: Test 7 – Non-linear potential for $\alpha = 1$, seed is 47. The transient state is depicted for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 1 compared to one with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 0.1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

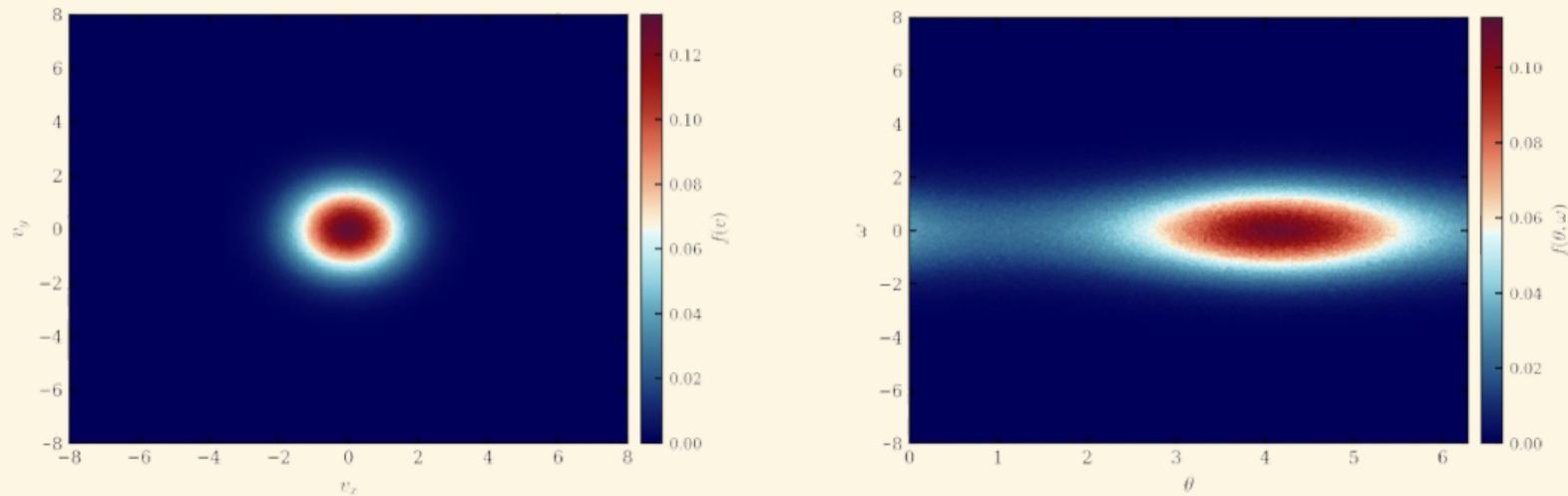


Figure: Test 7 – Non-linear potential for $\alpha = 1$, seed is 47. The Maxwellian steady state is reached for a simulation with $\Delta t = 0.01$, 10^7 particles, 256 bins and collision rate 1.

AN EXAMPLE: ROD-LIKE MOLECULES DSMC SIMULATIONS

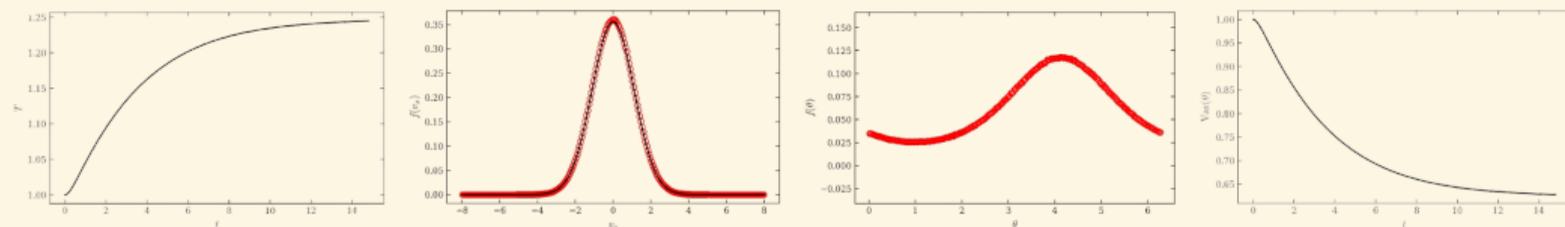


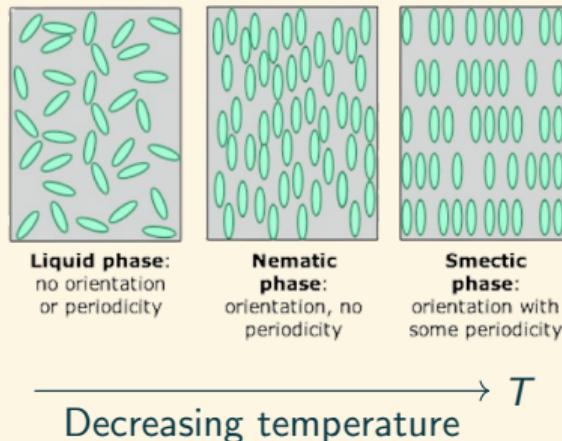
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ONSAGER STERIC POTENTIAL

Lastly we consider the Onsager potential for rod-like molecules, i.e.

$$\mathcal{W}(\theta_1, \theta_2) = L^2 |\sin(\theta_1 - \theta_2)|$$

where L is the length of the molecule. This potential can be derived considering purely steric interactions between two rod-like molecules, and has been used by Onsager in his seminal work on liquid crystals. In fact, this potential favours alignment between molecules via a purely entropic driven mechanism.



$$\mathcal{V}(\theta_1, t) = L^2 \int_0^{2\pi} \int_{-\infty}^{\infty} \partial_{\theta_1} (|\sin(\theta_1 - \theta_2)|) f(v_2, \theta_2, \omega_2, t) dv_2 d\theta_2 d\omega_2,$$

AN EXAMPLE: NEMATIC LIQUID CRYSTALS DSMC SIMULATIONS

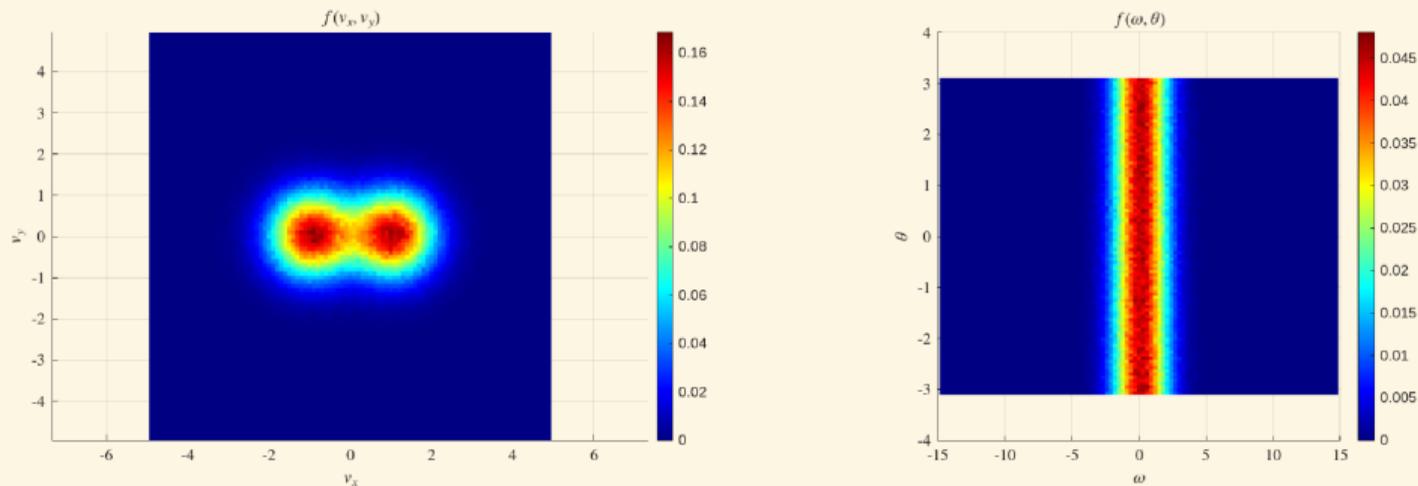


Figure: Test 4 - The Onsager potential for rod-like molecules. The grey dashed lines are the initial data, the black circles the numerical results, the solid red lines the steady state from Test 1. We observe a good accordance between numerical and exact results.

AN EXAMPLE: NEMATIC LIQUID CRYSTALS DSMC SIMULATIONS

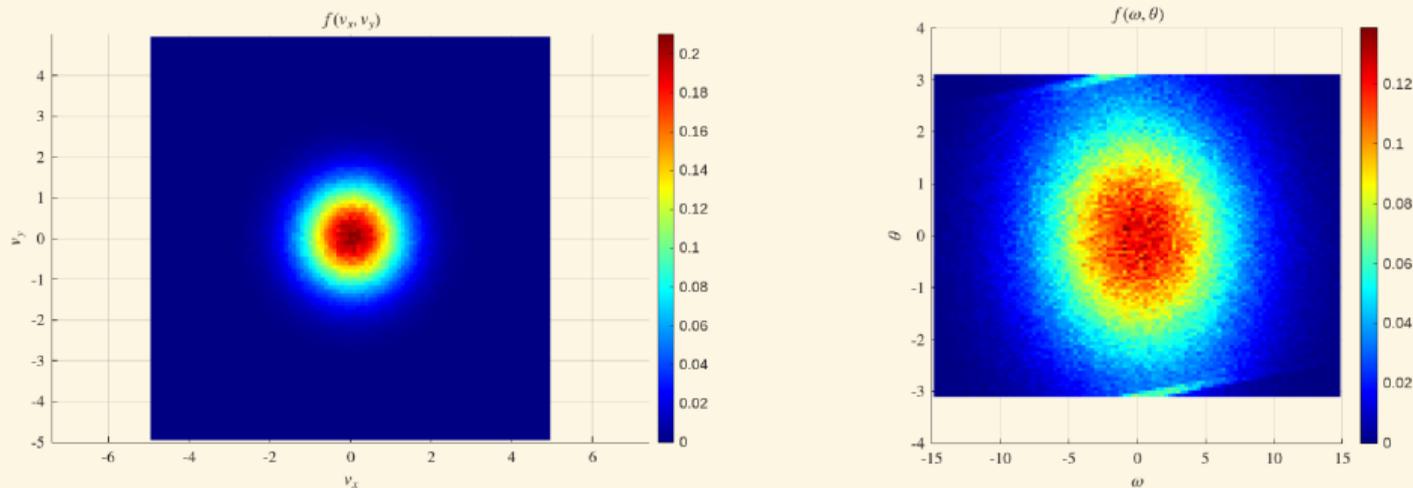


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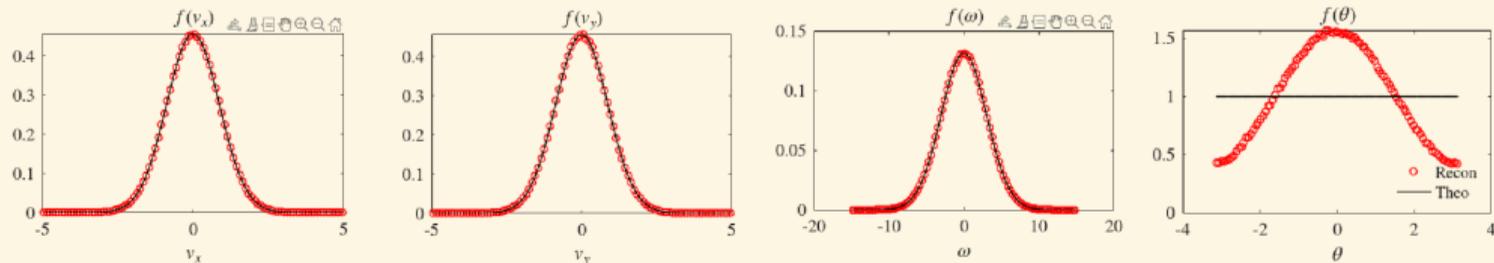


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THANK YOU!

New Challenges in the Kinetic Theory of Complex Fluids

UMBERTO ZERBINATI*, JOINT WORK WITH: J. A. CARILLO* ,P. E. FARRELL*,
A. MEDAGLIA*, G. RUSSO†